

# Solving Magnetospheric Acceleration, Reconnection and Turbulence (SMART)

#### **MISSION DEFINITION**

MMS-SMART will perform the first definitive study of magnetic reconnection in space and will test critical hypotheses about reconnection derived from the latest theories and spacecraft measurements. The SMART payload is optimized to make the required high-time-resolution measurements of plasmas, electric fields, and magnetic fields. SMART will also provide important supporting measurements of energetic particles and ion composition. The SMART plasma composition instrument is a new design that for the first time will solve the problem of identifying minor ions within regions of high proton fluxes. SMART will also solve known difficulties associated with low-energy plasma and electric-field measurements by including a flight-proven charge neutralization device. The accuracy of the electric field measurements will be further optimized by the inclusion of an electron-drift electric-field detector, which is immune to the plasma sheath effects that typically bedevil double-probe detectors. A three-phase low-inclination orbit strategy will probe the most likely reconnection sites on both the dayside magnetopause and in the magnetotail. Data from the three mission phases will be assimilated into continuously refined reconnection models in a way that will lead to fundamental advances in our understanding of reconnection. SMART will also address the MMS secondary science objectives.

## **SCIENCE OBJECTIVES**

Overall Scientific Objectives: To discover the detailed physics of the reconnection process including the factors that control it, its spatial distribution, and its temporal behavior.

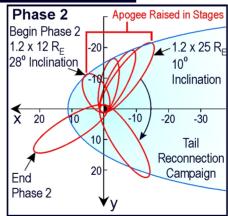
#### **Primary Science Questions:**

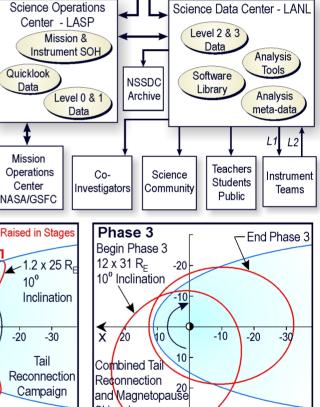
- 1) What are the kinetic processes responsible for collisionless magnetic reconnection, and how is reconnection initiated?
- 2) Where does reconnection occur in the magnetopause and in the magnetotail, and what influences where it occurs?
- 3) How does reconnection vary with time, and what factors influence its temporal behavior?
- 4) How do flux transfer events and plasmoids/magnetotail flux ropes form, and how do they evolve?

## Secondary Science Questions:

- 1) What is the role of inductive electric fields and wave-particle interactions in acceleration processes?
- 2) What are the properties and processes associated with magnetospheric turbulence?

#### Phase 1 6am End -20 Launch & Boost Phase 1 1.2 x 12 R<sub>E</sub> -10 28º Inclination midnight 20 -30 -20 **End Commissioning** Begin Science Phase 1 20 **GSE Coordinates**





Skimming

Campaign

MISSION OPERATIONS AND DATA FLOW

SMART Mission Project Office - SwRI

### **EDUCATION AND PUBLIC OUTREACH**

The MMS-SMART PI, **Dr. James L. Burch,** is personally committed to the SMART E/PO effort. The SMART E/PO lead is Rice University professor Patricia Reiff, a well-known leader in space physics E/PO and teacher enhancement programs. SMART E/PO products will be distributed through an established network of teacher organizations and museums as well as through SECEF and LWS outreach programs. SMART E/PO will

- Support three or four Master Teachers each year as part of Rice University's Master of Science Teaching Program.
- Develop two inquiry-based modules "Micro to Macro" and "What Changed?" aligned with national math and science standards for use in teacher workshops.
- Produce "Sounds of Space," which uses data "sonification" techniques to make SMART science accessible to visually
- impaired learners. Disseminate SMART science to students, teachers, and the public through the SMART web site, museum exhibits,
- immersive planetarium shows, and an updated version of the Space Weather CD-ROM. Evaluate during Phase B the addition of a small student-developed SSD to the SMART payload.

SCIENCE PAYLOAD					TEAMING ARRANGEMENT	
Fields *Spin Plane Double Probe (SDP; 4 ea.)					Southwest Research Institute: SMART Team Lead Primary Contractor Ion Composition Lead	
*Axial Double Probe (ADP; 2 ea.) *Electron Drift Instrument (EDI; 2 ea.)					Central Instrument Data Processor Lead	
*Analog Flux Gate Magnetometer (AFG)					Payload Integration Lead	
Digital Flux Gate Magnetometer (DFG) Search Coil Magnetometer (SCM)					University of New Hampshire: Fields Suite Lead	
*Central Electronics Box (CEB)  Fast Plasma					Goddard Space Flight Center: Fast Plasma Lead Theory & Modeling Lead	
*Dual Electron Sensors (DES; 4 ea.)					, ,	
*Dual Ion Spectrometers (DIS; 4 ea.) *Instrument Data Processing Unit (IDPU)					Applied Physics Laboratory: Energetic Particles Lead  Austrian Academy of Sciences: ASPOC Lead	
Plasma Composition					•	
Hot Plasma Composition Analyzer (HPCA)  Energetic Particles					Other Hardware-Providing Institutions: Fields	
*Fly's Eye Energetic Particle Sensor (FEEPS; 2 ea.) Energetic Ion Spectrometer (EIS)					Austrian Academy of Sciences French Center for Terrestrial and Planetary Environments	
Spacecraft Potential Control  *Active Spacecraft Potential Control (ASPOC; 2 ea.)				C; 2 ea.)	Swedish Institute of Space Physics Swedish Royal Institute of Technology	
Payload Processor					Technical University of Braunschweig University of California at Los Angeles	
*Central Instrument Data Processor (CIDP, redundant)  (*Element of minimum payload)					University of Colorado at Boulder	
PAYLOAD LAYOUT					University of Iowa	
ASPOC (2x) FIELDS CEB (1x)					Fast Plasma Investigation Japan Aerospace Exploration Agency	
ADP.					Ion Composition	
(2x) FEEPS				FEEPS (2x)	Applied Physics Laboratory Lockheed Martin Advanced Technology Center	
				and the second second	Los Alamos National Laboratory	
FPI/DIS (4x)					Energetic Particles	
				FDVDF0	The Aerospace Corporation  MISSION MANAGEMENT	
(4x)				FPI/DES (4x)	Principal Investigator	
SDP					Dr. James L. Burch, SwRI	
(4x)					Project Manager Mr. Bill Gibson, SwRI	
HPCA + CIDP (1x)					Payload Systems Engineer	
(ZX)				(2x)	Mr. Ron Black, SwRI	
PAYLOAD MASS & POV					Fields Team Lead Prof. Roy Torbert, UNH	
Subsystem	Mass (kg)		Power (Watts)		Fast Plasma Lead	
Fields	CBE 32.40	Contingency 3.72	CBE 13.84	Contingency 1.89	Dr. Thomas E. Moore, GSFC  Plasma Composition Lead	
Fast Plasma	15.00	1.60	13.84	2.27	Dr. David Young, SwRI	
HPCA	5.88	1.17	6.13	1.23	Energetic Particle Lead	
Energetic Particles	2.54	0.45	1.52	0.21	Dr. Barry Mauk, APL Theory and Modeling Team Lead	
ASPOC	3.68	0.37	1.96	0.49	Dr. Michael Hesse, GSFC	
CIDP	0.76	0.11	1.92	0.19	Education & Public Outreach Lead Prof. Patricia Reiff, Rice Univ.	
Misc.	4.79	1.19	N/A	N/A		
Total	65.05	8.61	36.71	6.28	SCHEDULE AND COST SUMMARY  1st Delivery	
SRR PDR CDR					CDR to S/C Launch	
Phase A Bridg		Bridge Phase	Ph	ase B	Phase C/D Phase E	
<b>10/03-09/04</b> \$970K FY03			<b>10/04-1/06</b> \$20,432K FY03		<b>02/06-04/10 04/10-04/13</b> \$88,300K FY03 \$12,526K FY03	
\$1000K RY \$22,010K RY			\$22,	010K RY	\$100,463K RY \$16,277K RY	
Total Cost \$139.75 M RY 20% cost reserves on Phase C/D, 10% cost reserves on Phase E						
	TA004133					